

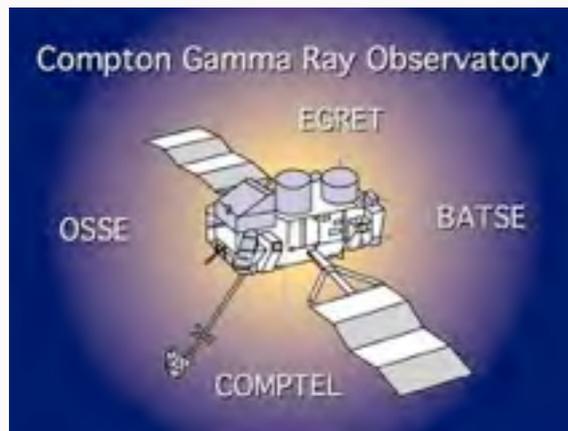
# TERRESTRIAL GAMMA-RAY FLASHES: FERMI'S EMERGING ROLE AS AN EXCELLENT STORM CATCHER



Valerie Connaughton

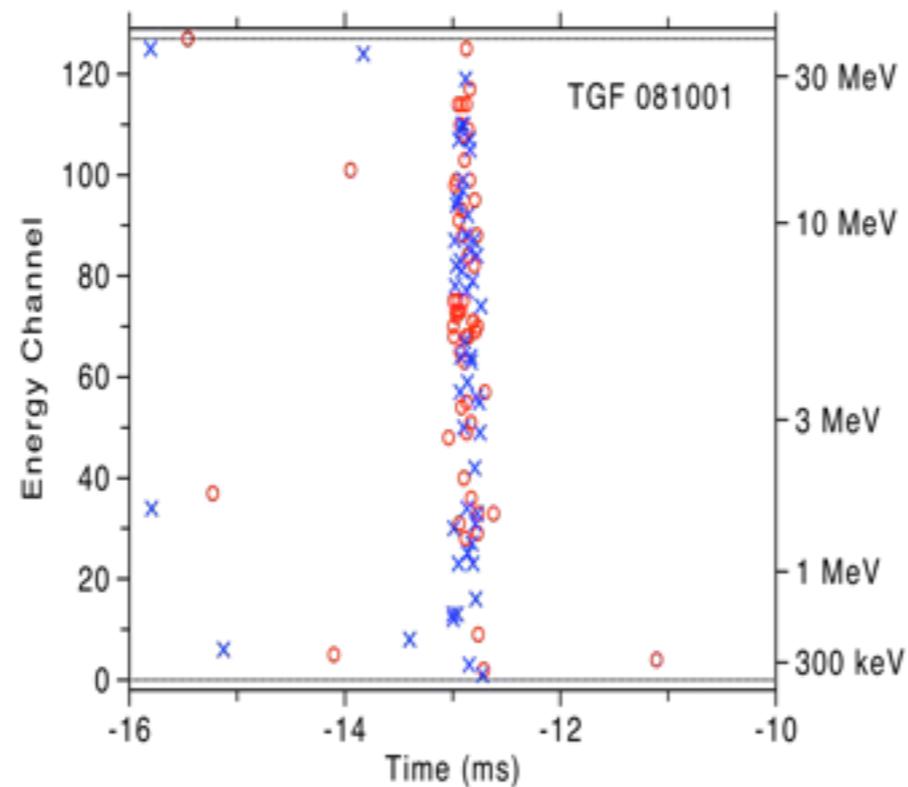


# TERRESTRIAL GAMMA-RAY FLASHES



Fishman et al. 1994

- Smith et al. 2005
- Grefenstette et al. 2009
- Gjesteland et al. 2012

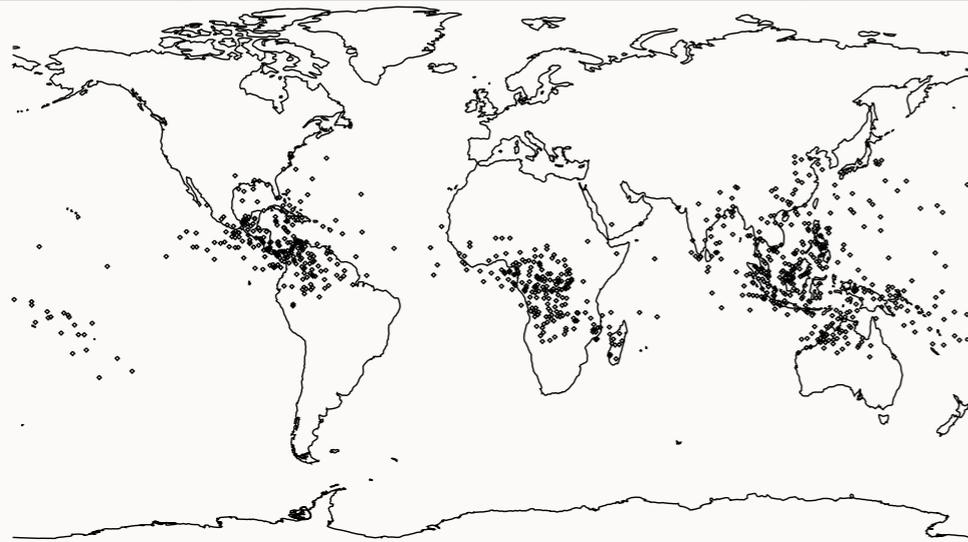


Fuschino et al. 2011

- Briggs et al. 2010
- Fishman et al. 2011

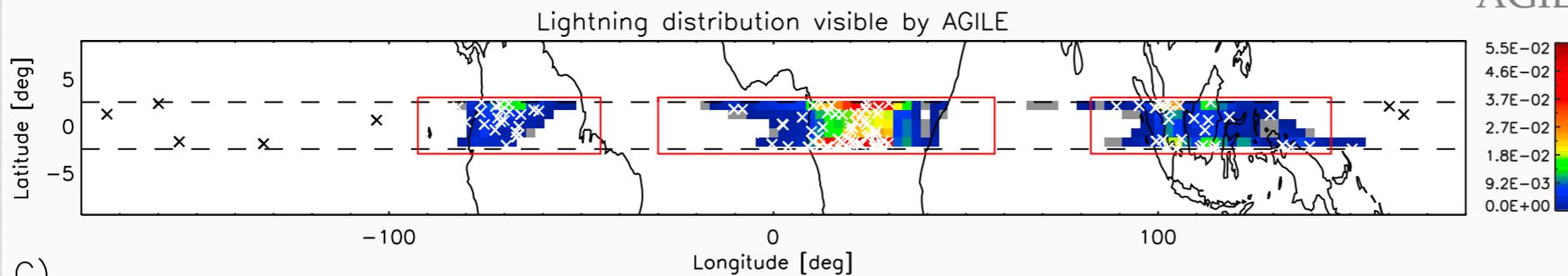


# SPACECRAFT NADIRS WHEN TGFS ARE DETECTED



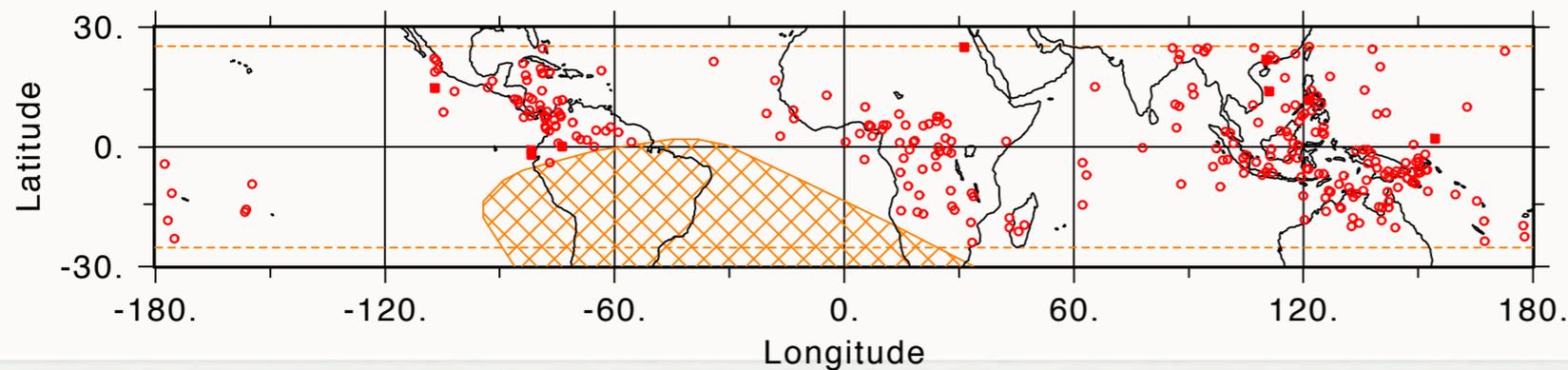
RHESSI: Grefenstette et al. 2009

AGILE: Fuschino et al. 2011



C)

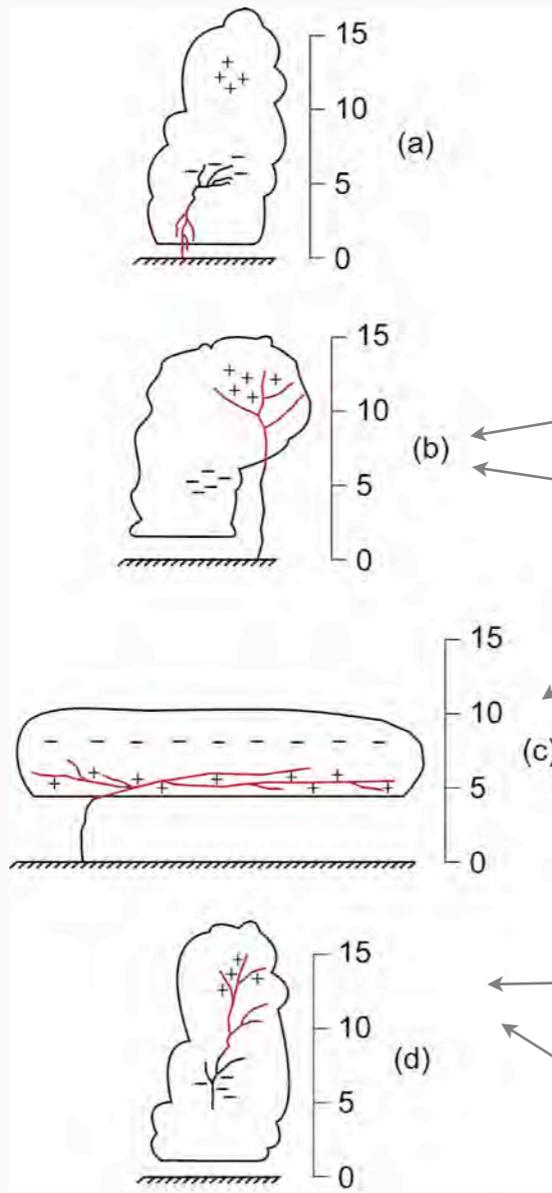
273 GBM TGFs



GBM Latest Map!  
See also Briggs et al. 2010,  
Fishman et al. 2011

# A SHORT LIGHTNING PRIMER

Radio detection of lightning through sferics.  
VLF 3 - 30 kHz, ELF 3 - 3000 Hz ULF < 3 Hz



Inan et al. 1996 (2 BATSE TGFs)  
Positive polarity

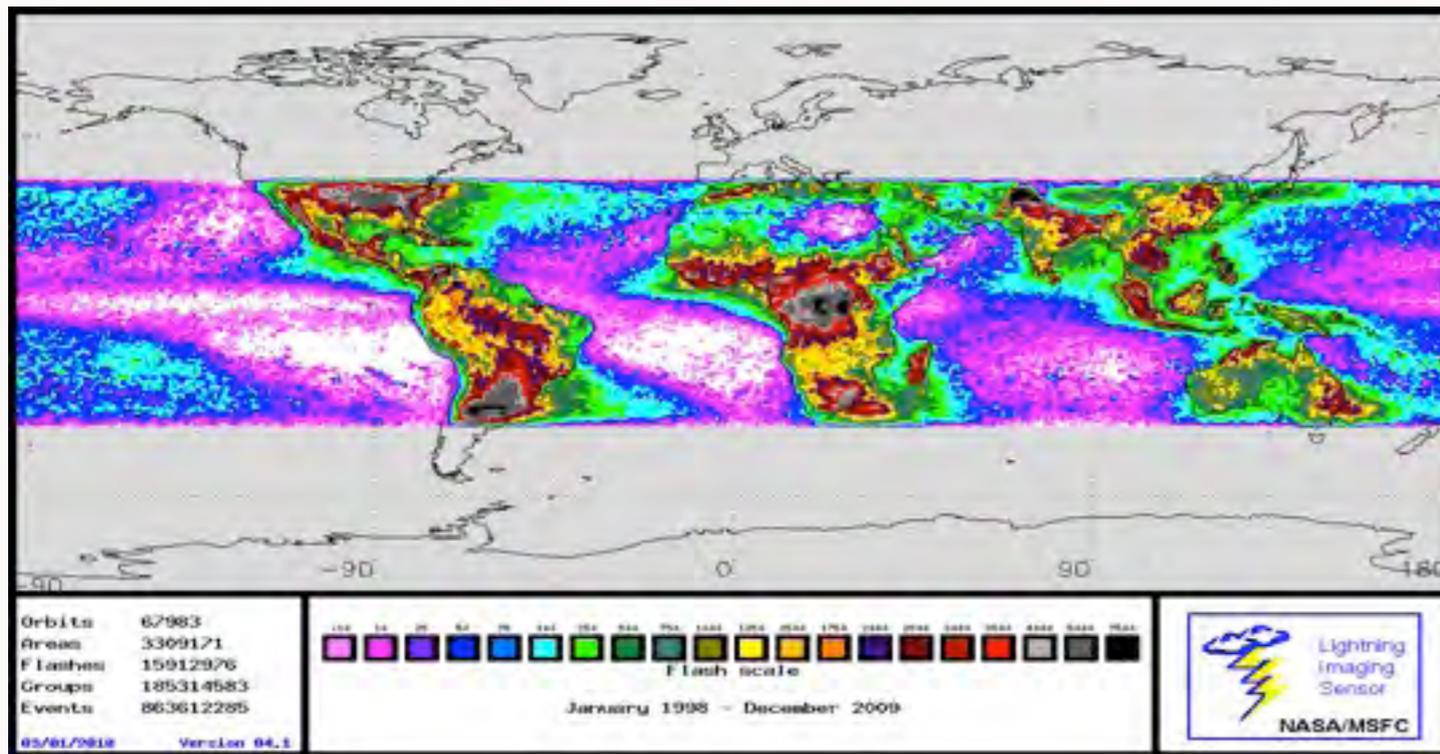
Cummer et al. 2005 (13 RHESSI TGFs)  
Average charge moment 49 C km

Dwyer and Smith 2005  
Source height 15-21 km  
(RHESSI spectrum)

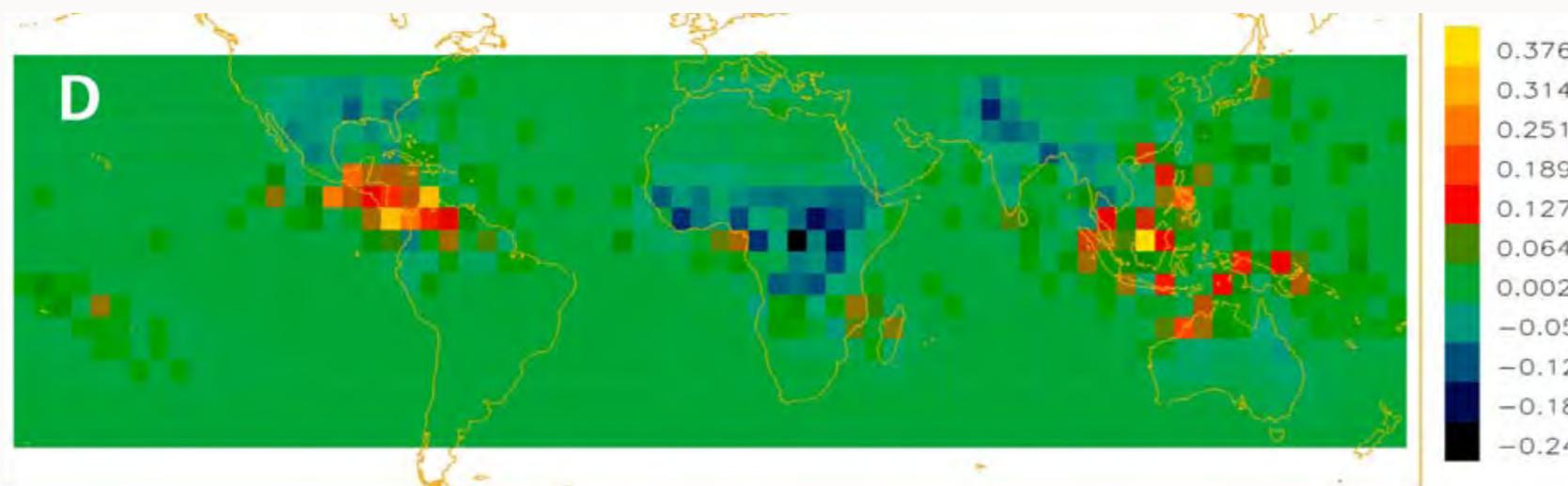
Stanley et al. 2006 (6 RHESSI TGFs)  
Source heights 11-13 km

Williams et al. 2006

# DO THE NADIRS FOLLOW THE GLOBAL LIGHTNING DISTRIBUTION?



LIS-OTD map

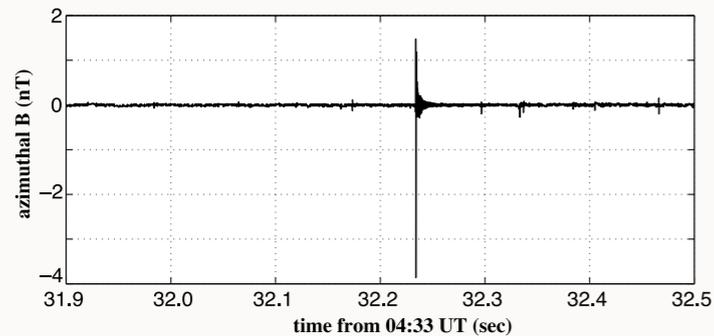


TGF excesses/deficits relative to LIS/OTD

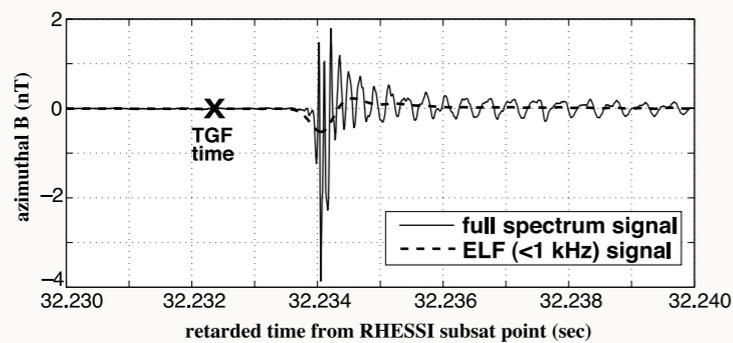
Smith et al. 2010

Not a one-to-one correspondence: instrumental effects, or meteorology affects TGF-production (Splitt et al. 2010)

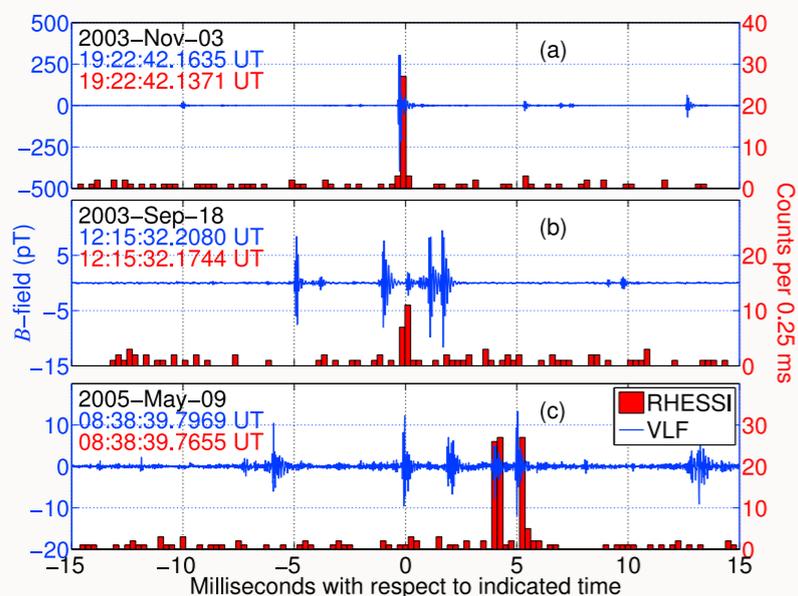
# TGF CORRELATION WITH INDIVIDUAL LIGHTNING DISCHARGES: CHICKEN OR EGG?



Cummer et al. 2005 (13 RHESSI TGFs)  
from -0.7 to 3 ms with average -1.24 ms



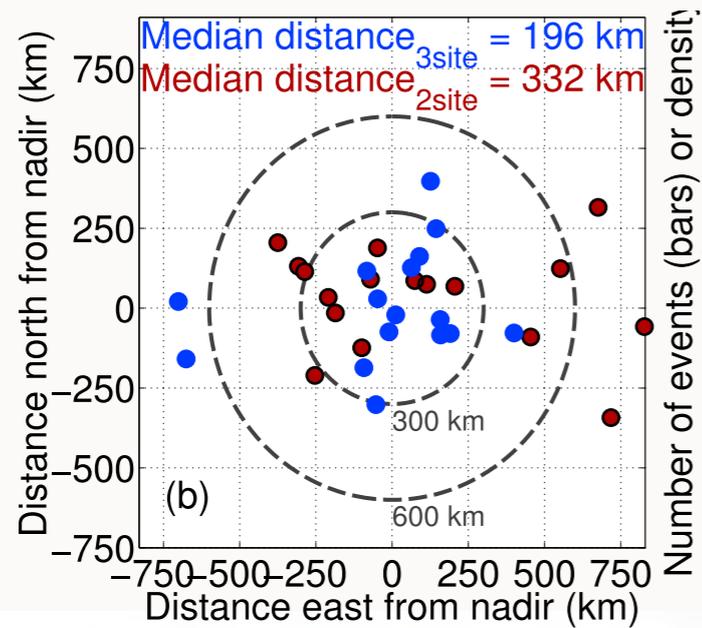
Inan et al. 2006 (88 RHESSI TGFs)  
average  $+0.879 \pm 2.41$



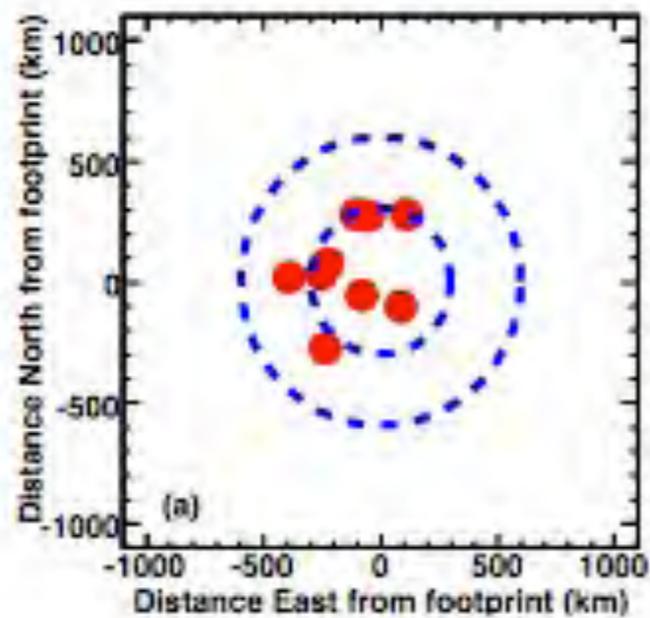
Cohen et al. 2010 (36 RHESSI TGFs)  
 $-0.5 \text{ ms} \pm 2.4 \text{ ms}$

RHESSI clock is uncertain to  $\sim 2 \text{ ms}$

# GEOLOCATION OF TGFS USING RADIO NETWORKS

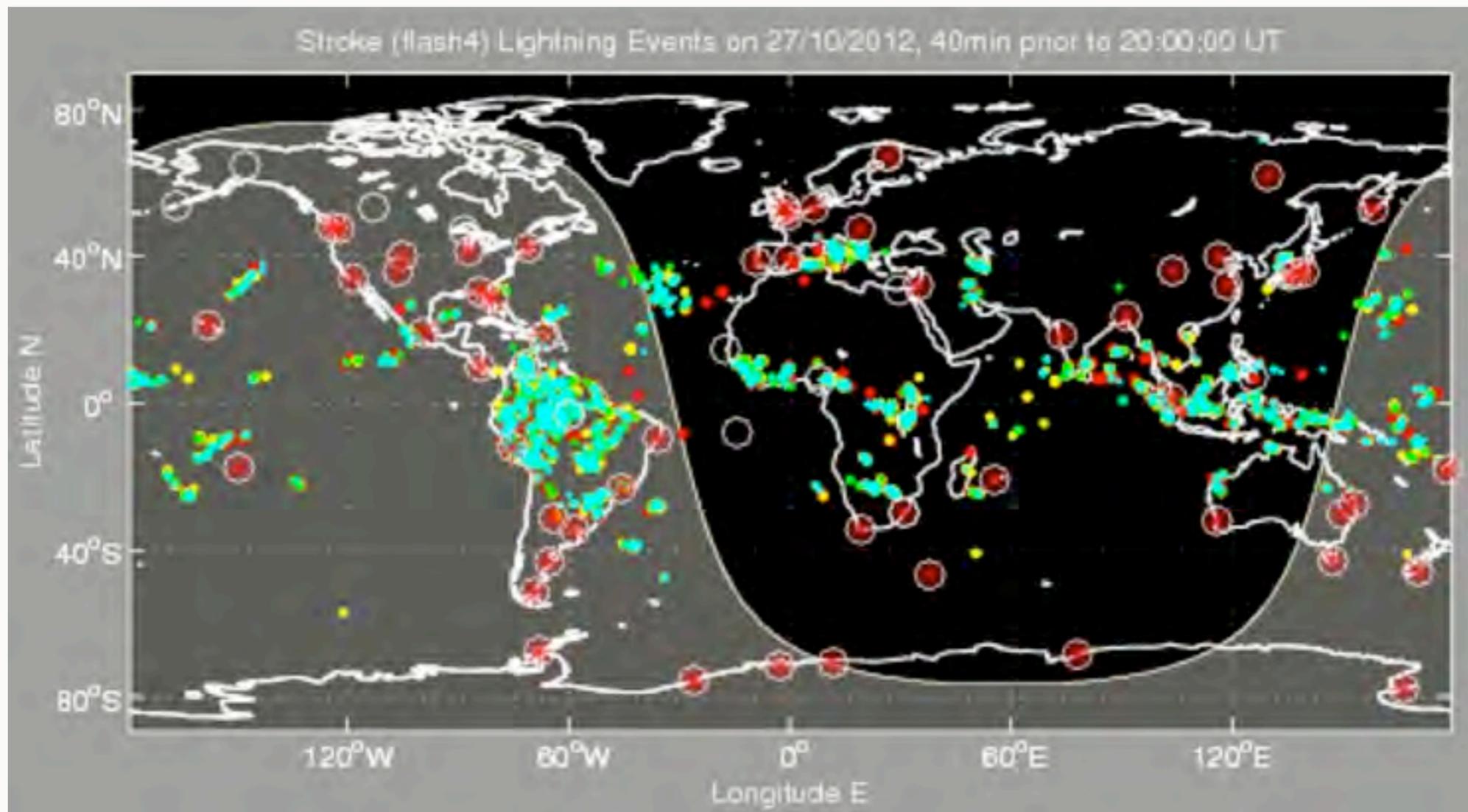


Cohen et al. 2010 (36 RHESSI TGFs)  
2-3 pairs of antennae



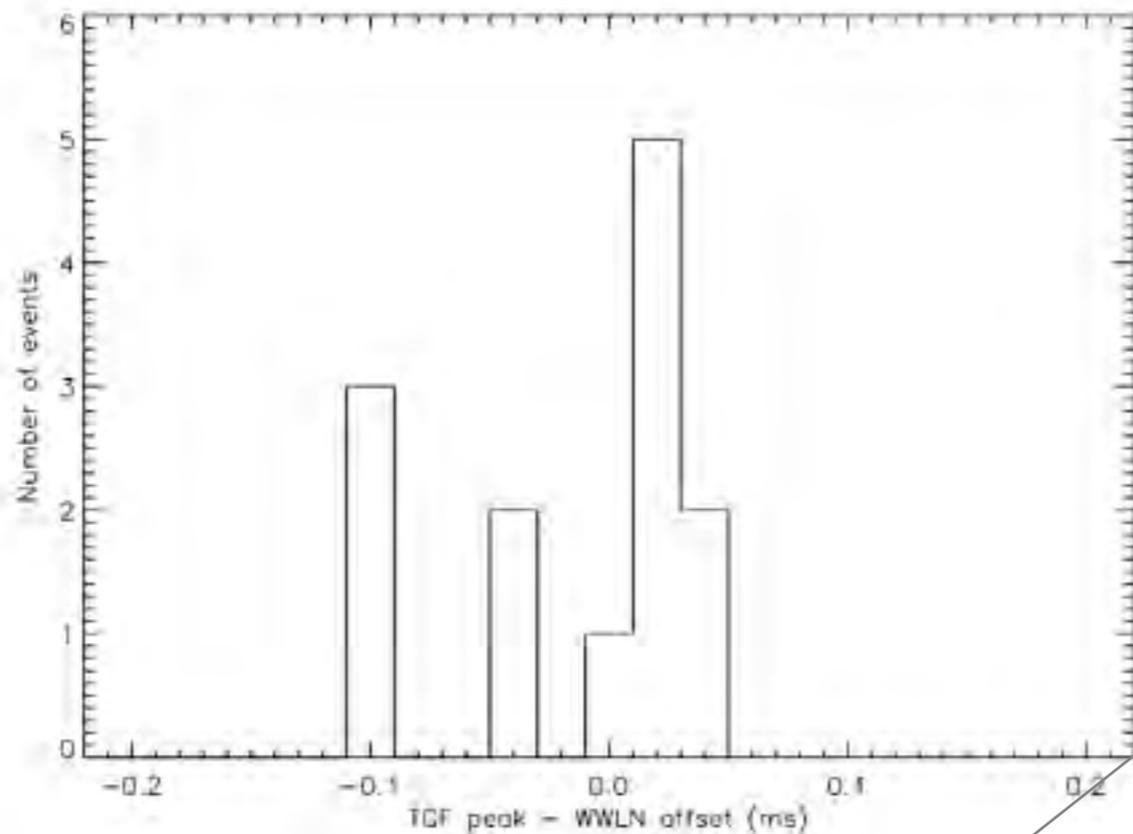
Geolocation can also be  
done using gamma rays:  
AGILE (Marisaldi et al.  
2010)

# WORLD WIDE LIGHTNING LOCATION NETWORK (WWLLN)



<http://webflash.ess.washington.edu/>

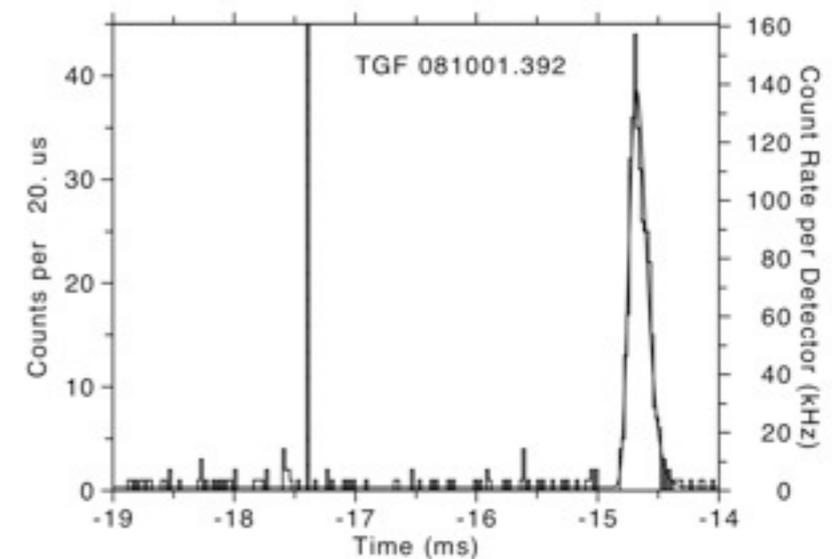
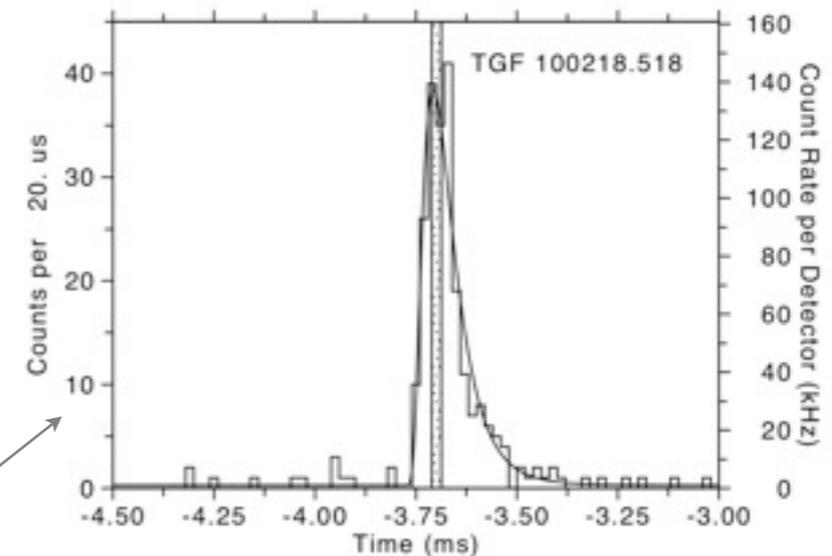
# GBM TGF PEAK TEMPORAL OFFSETS FROM WWLLN DISCHARGES FOR 15/50 ASSOCIATIONS (5 MS, 300 KM)



13/15 WWLLN discharges simultaneous with TGF peak.

... but 2 are ms away, either side of TGF pulse

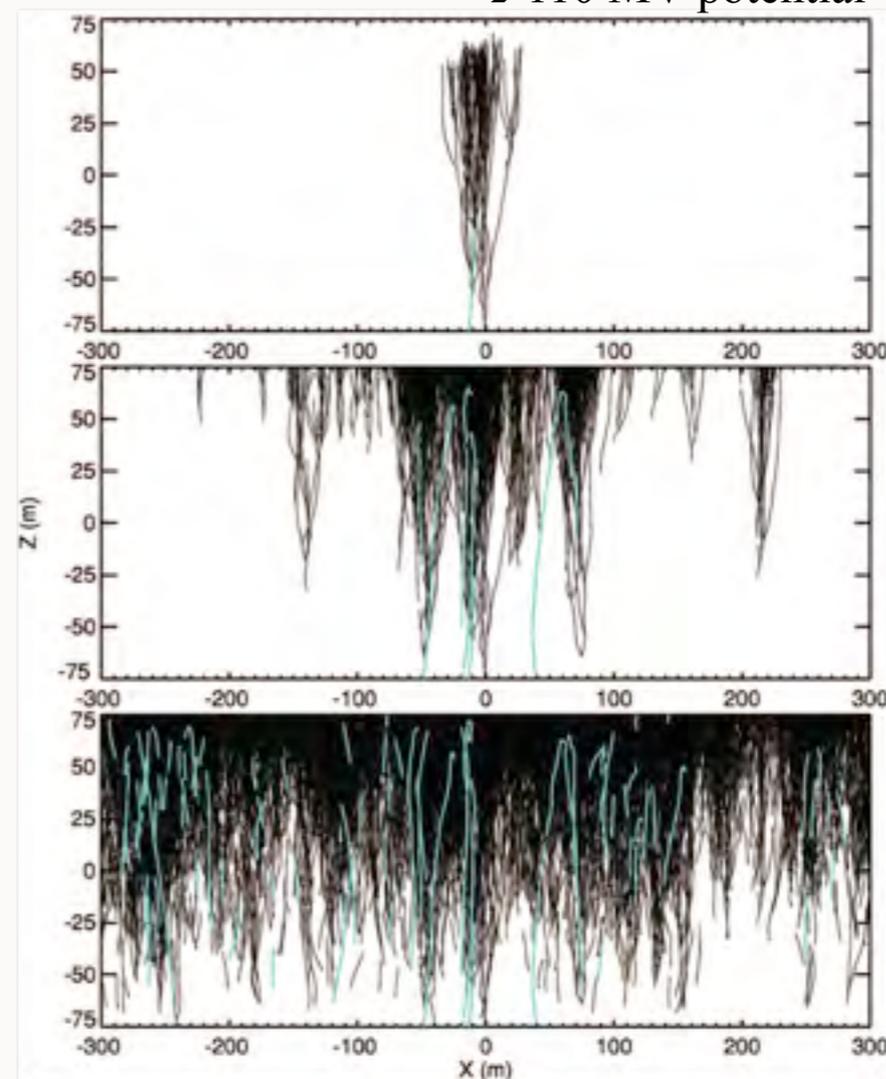
VC et al. (2010)



# BREMSSTRAHLUNG OF ELECTRONS ACCELERATED IN ELECTRIC FIELDS REPRESENTS WELL THE OBSERVED GAMMA-RAY EMISSION

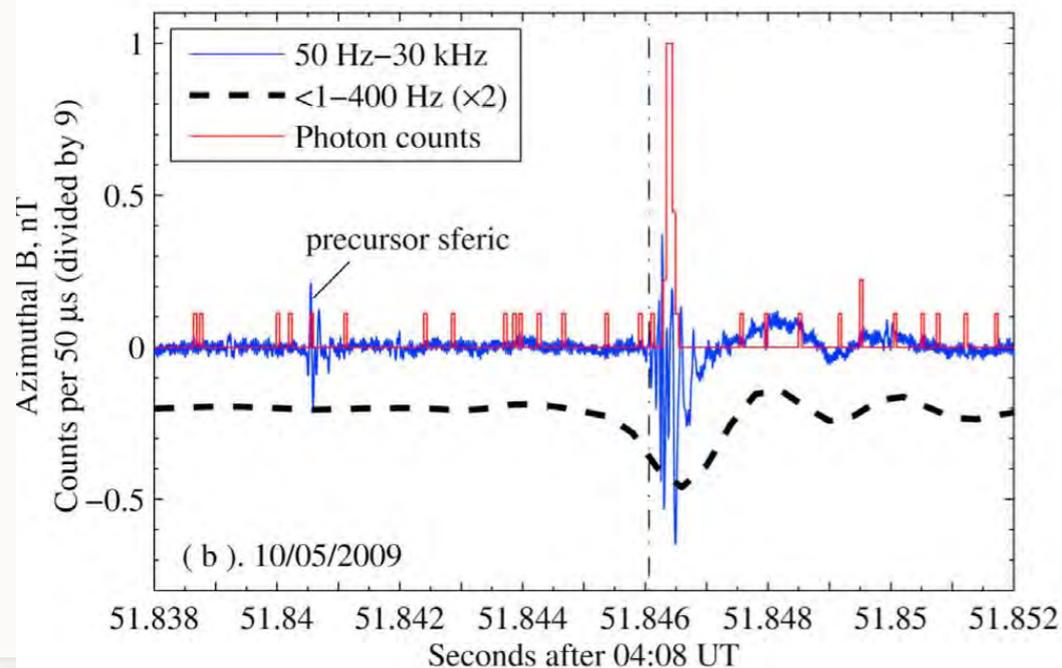
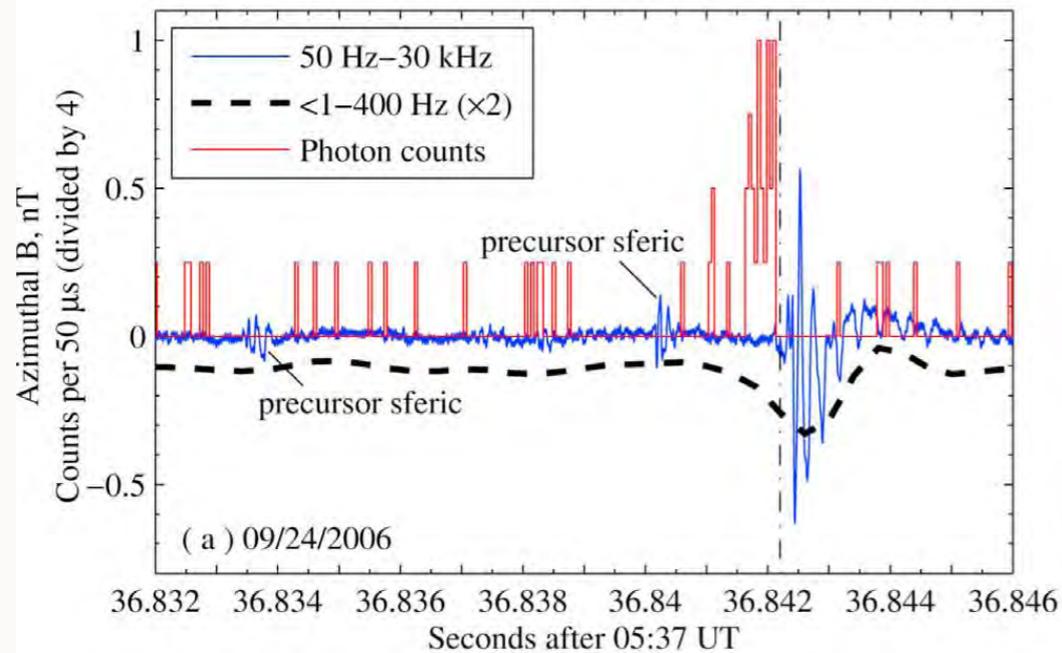
- Runaway Relativistic Electron Avalanche (Gurevich et al 92)
- Source of electrons? Need  $10^{16}$  photons, or  $10^{17}$  electrons, with  $E > 20$  MeV for observed flux
- Source of electric field and potential difference?
  - Thunderstorm ambient field - feedback produces more e's (Dwyer et al. 2007)
  - Alternative: Lightning leader tips - large potentials in tips. (Carlson et al. 2009, Celestin & Pasko 2011)

$E = 750$  kV / m  
for 150 m,  
→ 110 MV potential



Close temporal relation expected in leader tip model, not feedback model.

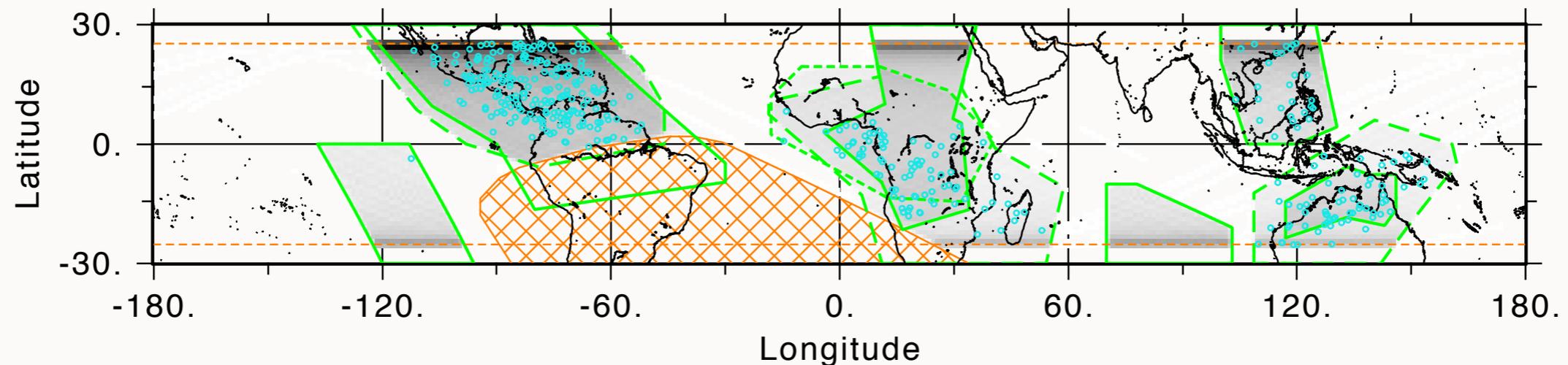
# 2 TYPES OF VLF RADIO PULSES ASSOCIATED WITH TGFS?



Lu et al (2010) 9 RHESSI TGFS

VLF and ULF pulses  
Slow ULF pulse with VLF sferics  
Fast VLF sferics occur too  
Fast VLF sferics can occur ms away

# GBM OFFLINE SEARCH



TTE event data gathered over select parts of orbit (moving boxes)

>1000 hours gathered

TGFs sought on-ground

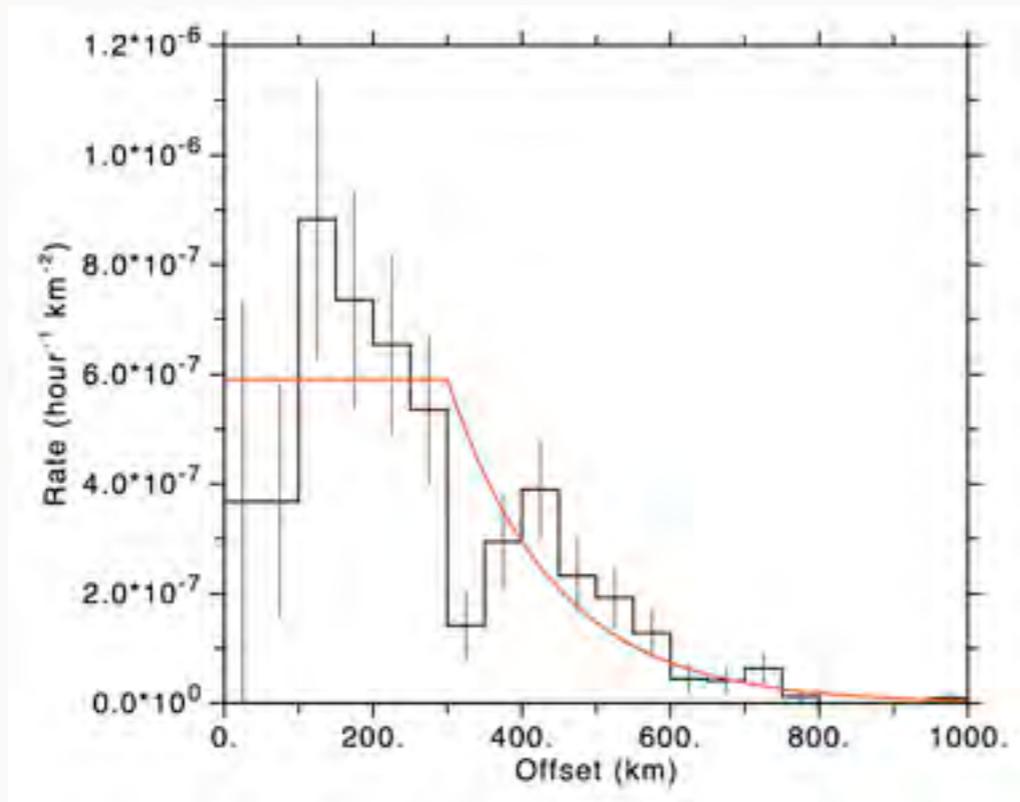
Cosmic ray rejection critical factor

TGF rate 10x triggered rate in same region. 384 offline only vs 39 onboard.

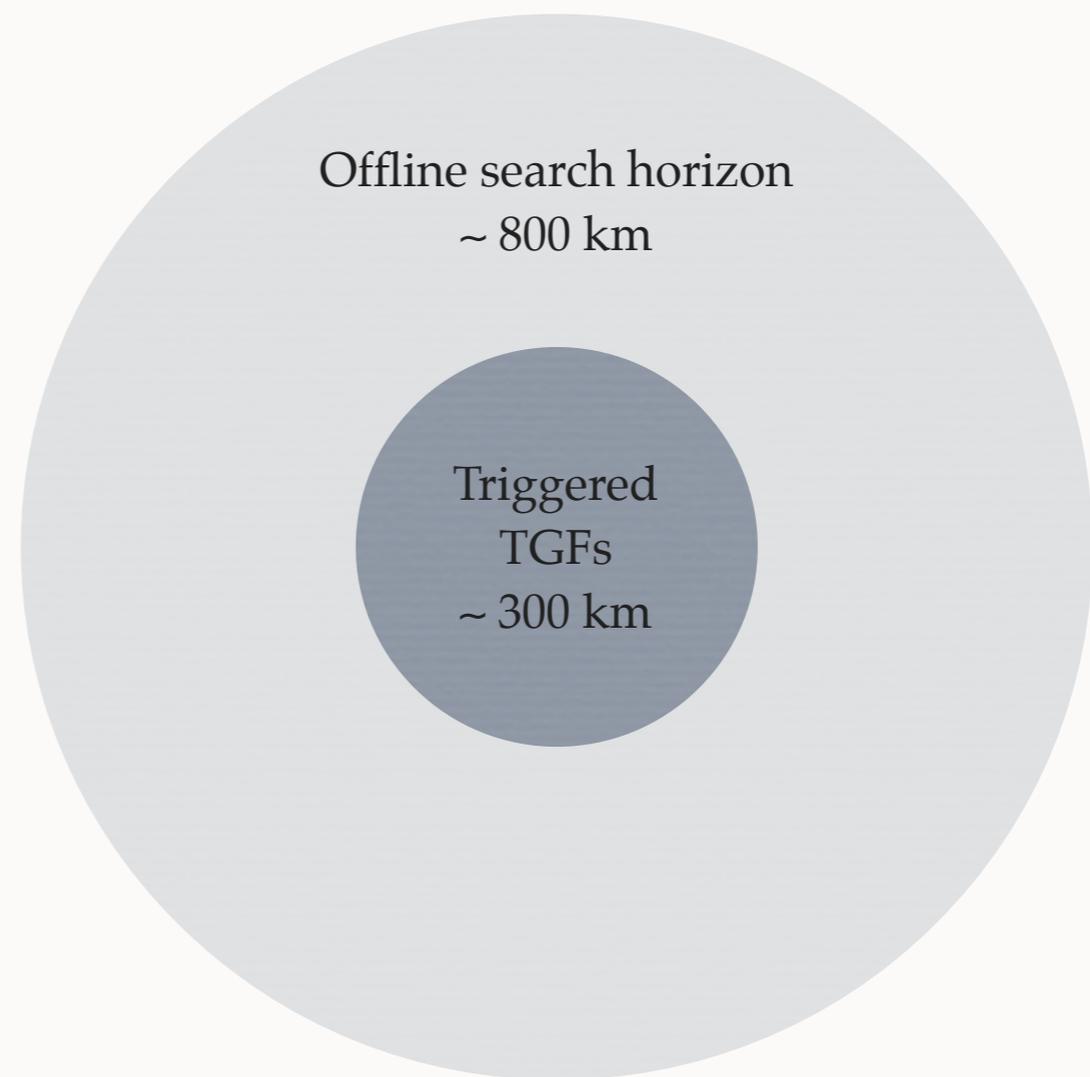
Briggs et al. (2012)

Shaolin Xiong Tuesday Instrument session

# WHAT IS GBM'S FOOTPRINT FOR TGF DETECTION?

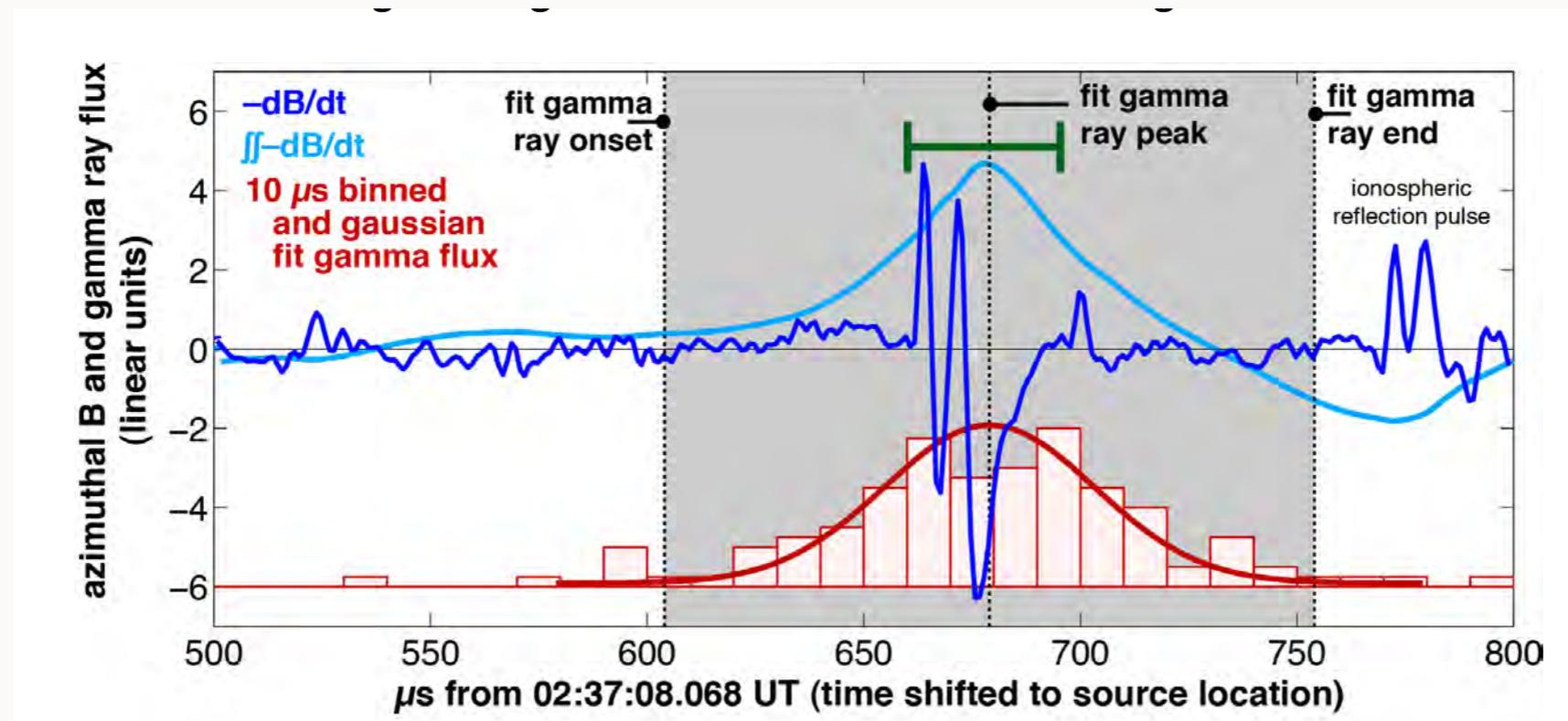


Briggs et al. (2012)



# MORE ON THE TWO TYPES OF RADIO PULSES ASSOCIATED WITH TGFS

- Using the Duke sensors in the US... now within reach of Fermi

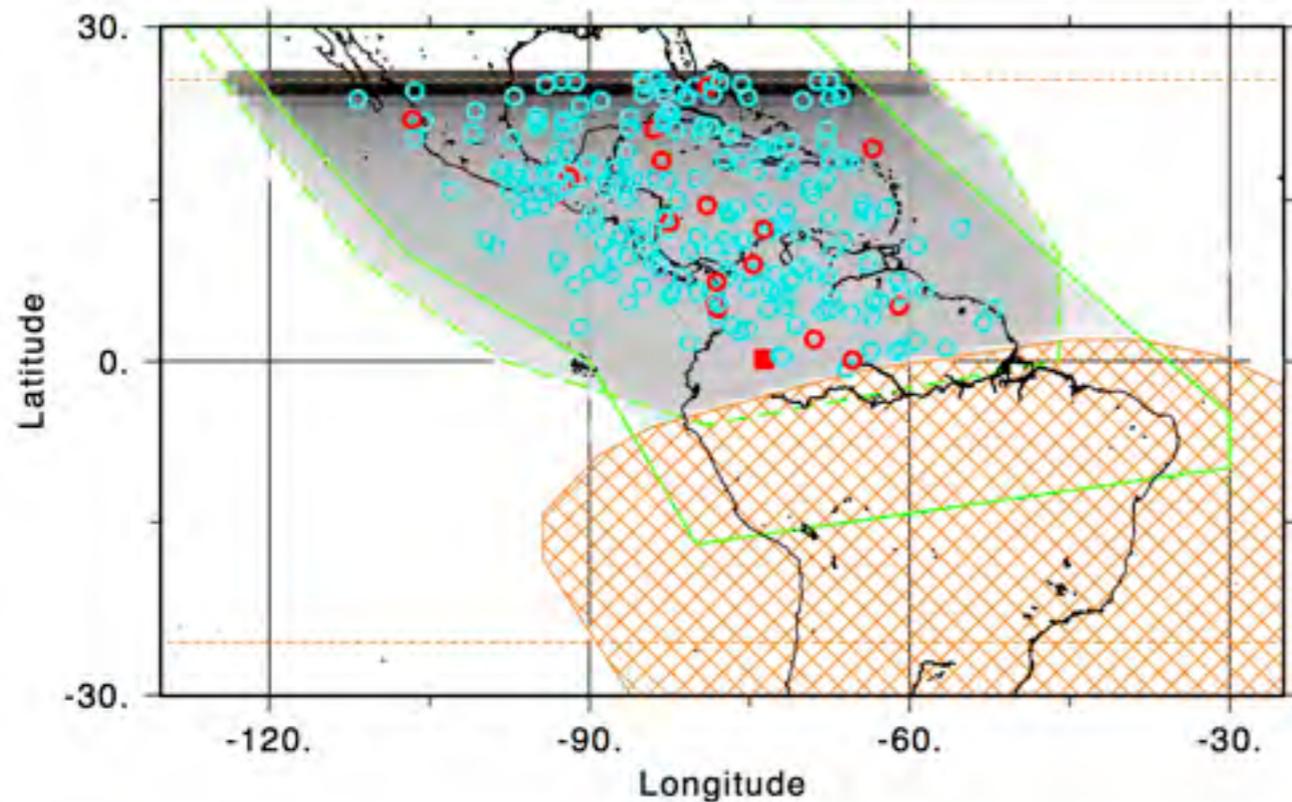


Could this slow varying component be current from the TGF itself?

Cummer et al. (2011)

Dwyer (2012)

# WHAT IS THE GLOBAL TGF RATE? DOES IT FOLLOW GLOBAL LIGHTNING RATE?



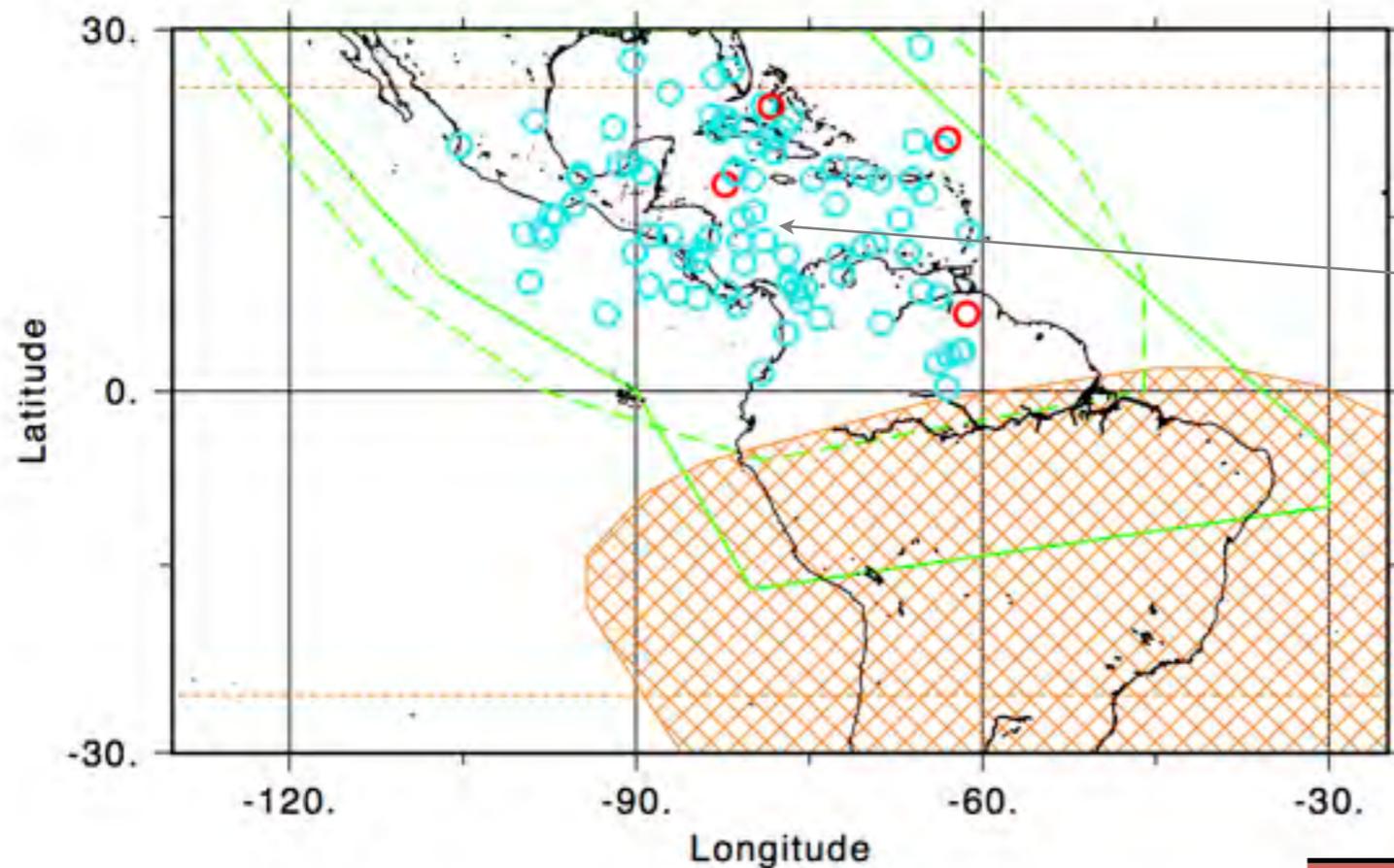
400,000 per year!

cf. Ostgaard et al. 2012  
 Smith et al. 2011  
 Fuschino et al. 2011  
 Carlson et al. 2011

Briggs et al. (2012)

Region	Ratio
Average	$(3.8 \pm 0.2) \times 10^{-4}$
Americas	$(4.9 \pm 0.3) \times 10^{-4}$
Africa	$(2.3 \pm 0.2) \times 10^{-4}$
Asia	$(2.7 \pm 0.4) \times 10^{-4}$
Australia	$(8.6 \pm 1.0) \times 10^{-4}$

# DISTRIBUTION OF TGF POSITIONS USING WWLLN GEOLOCATIONS

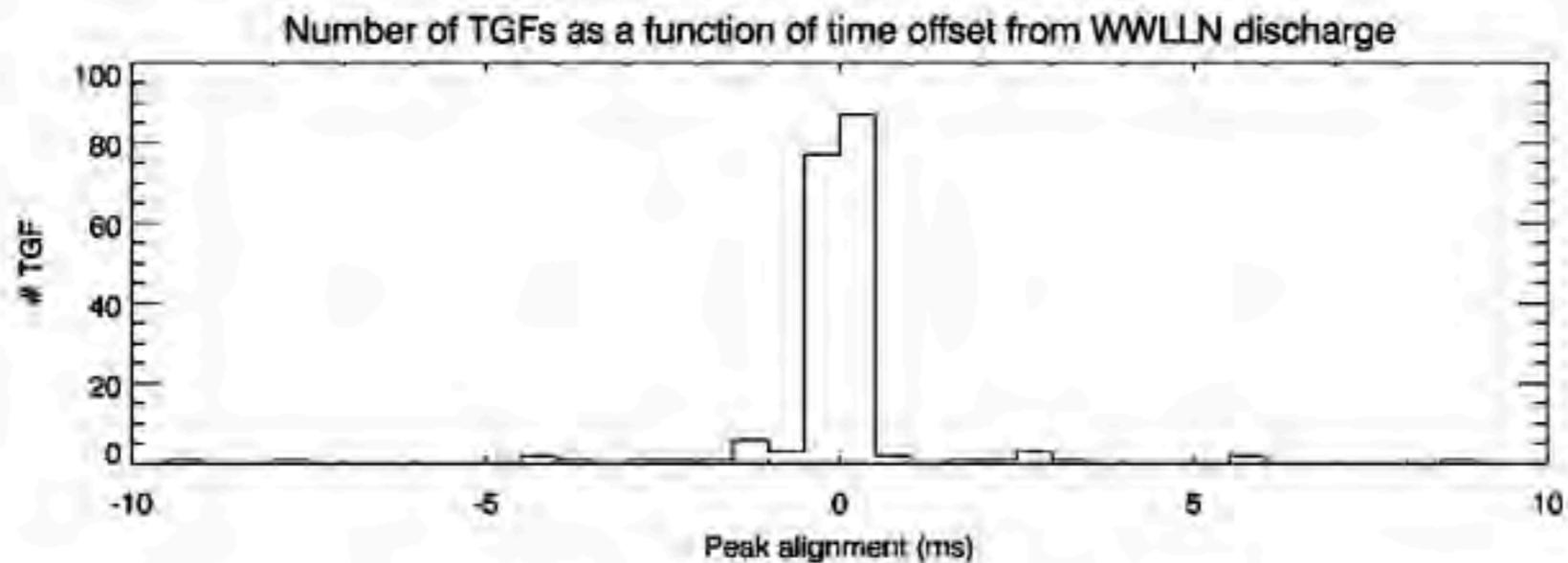


Coastal regions  
+  
Intertropical Convergence Zone  
(see Splitt et al. 2010)

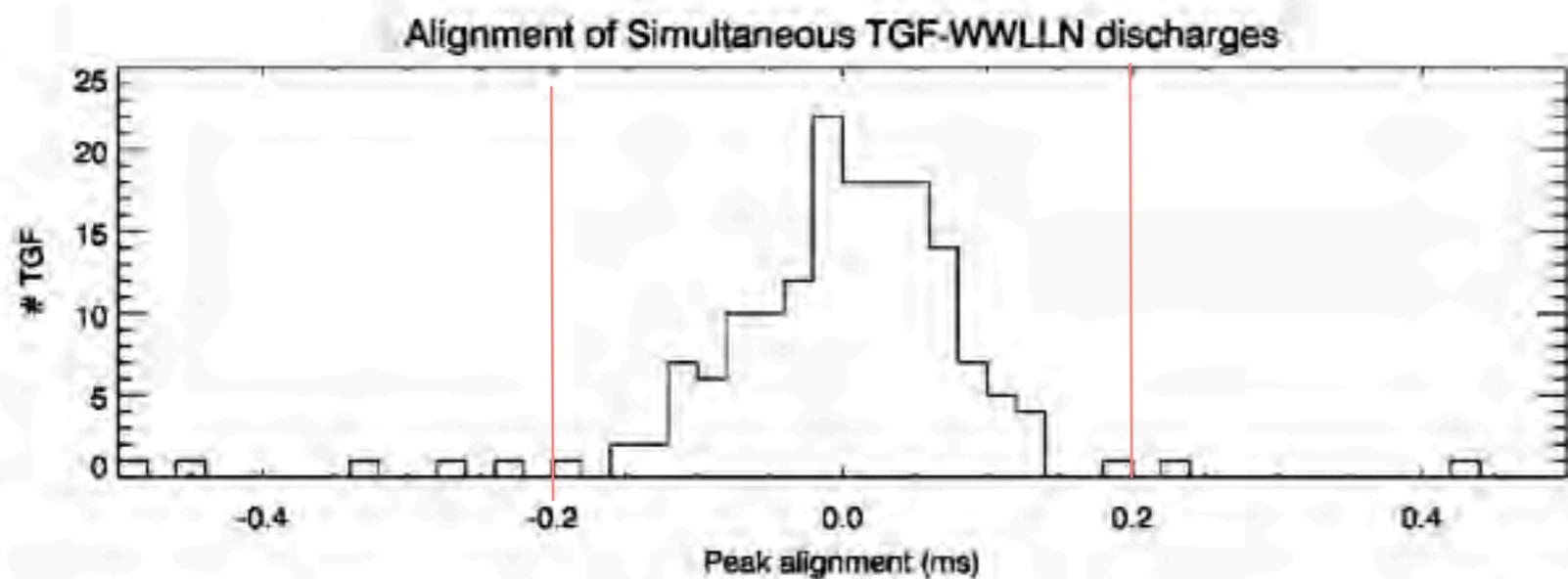
Briggs et al. (2012)

Region	Ratio
Average	$(3.8 \pm 0.2) \times 10^{-4}$
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# NEW GBM TGF-WWLLN DISCHARGE OFFSET DISTRIBUTION



186/594 TGFs  
have WWLLN  
associations.

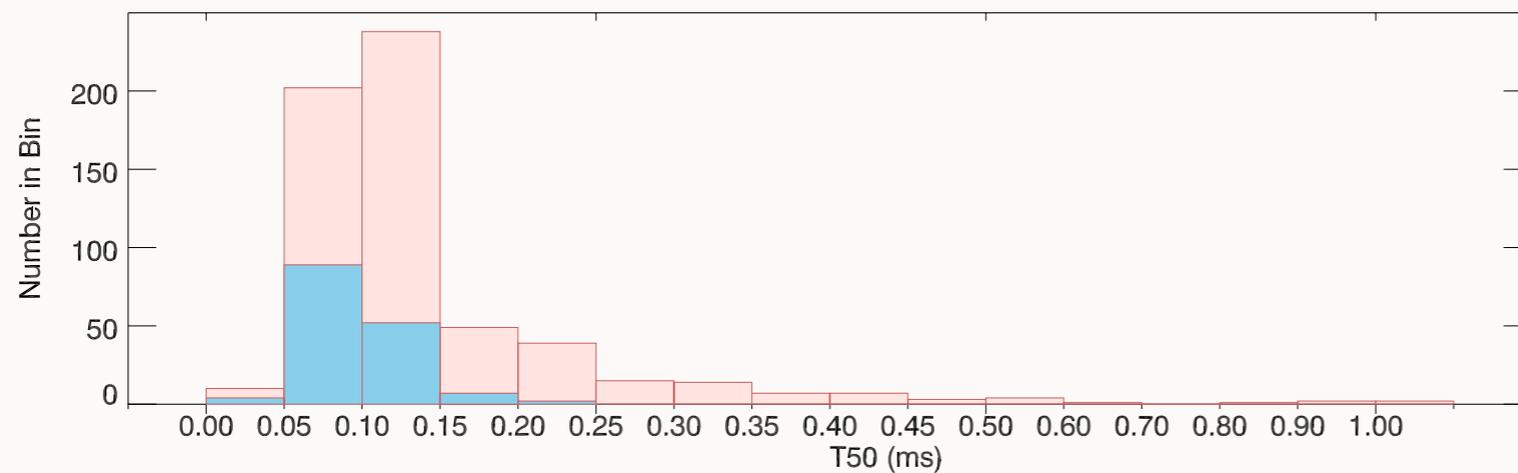


83% WWLLN  
associated discharges  
are simultaneous  
with TGF peak.

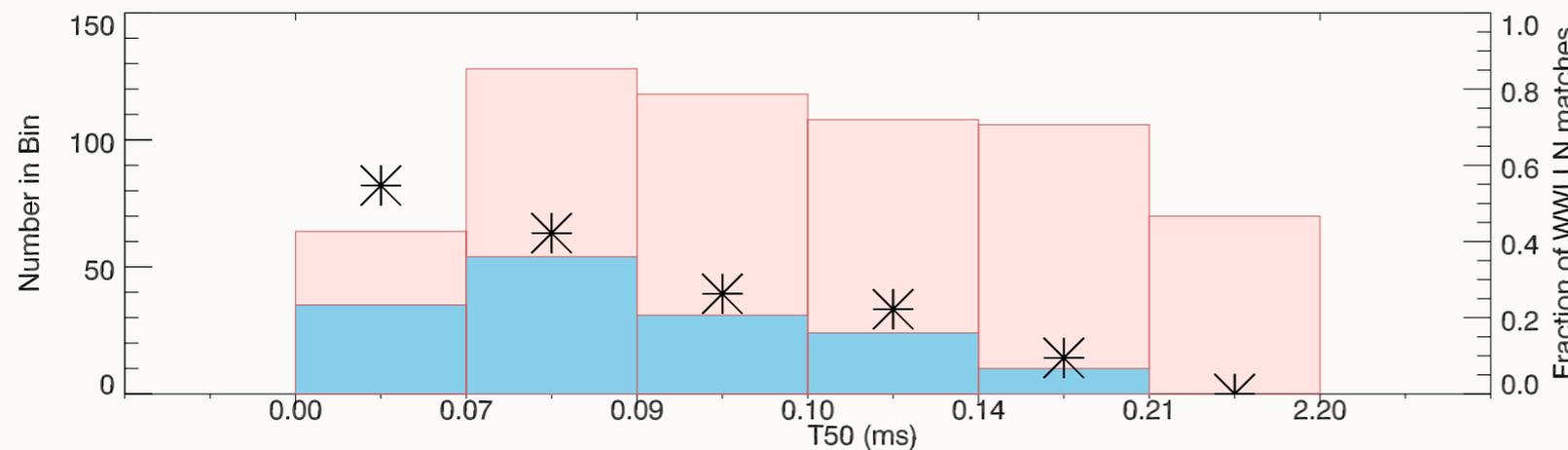
VC et al. (2012)

594 TGFs (triggered + offline)

154 TGFs with a simultaneous WWLLN discharge



Duration (T50) Distribution



WWLLN Match rate  
as function of  
TGF duration (T50)

VC et al. (2012)

# VLF SIGNALS FROM TGFS

- Electrons in RREA generate current => for a given total charge, short TGFS will generate more current.

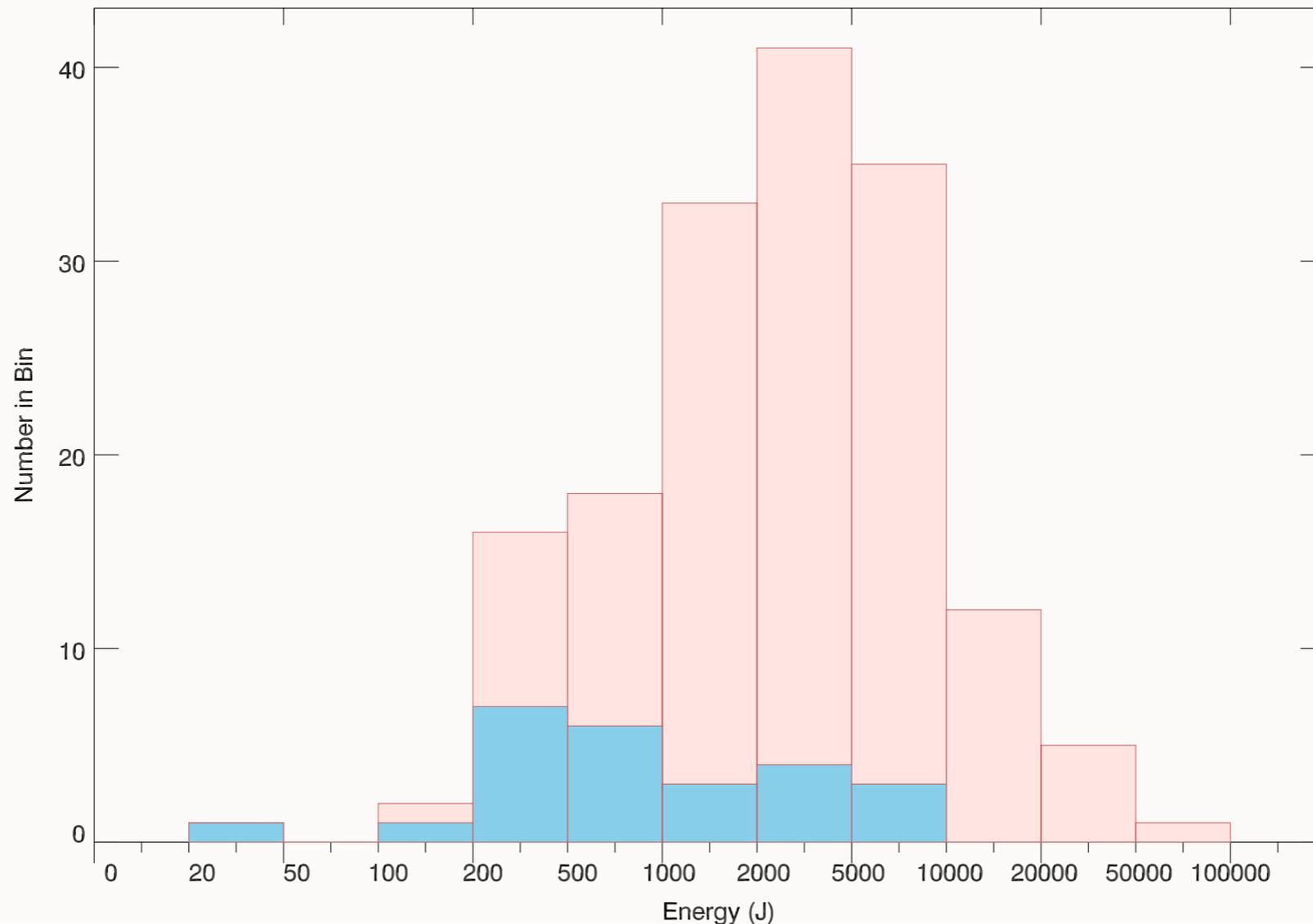
$$E(\omega) = -i\omega \frac{e\alpha\tau_a\mu_e EN_{re}\Delta z \sin\theta}{\sqrt{2\pi}4\pi\epsilon_0 c^2 R} \exp\left(\frac{-\omega^2(0.74T_{50})^2}{2}\right)$$

Eq 3 of VC et al. 2012

- Short TGFS are more likely to have a detectable WWLLN signal
- Non-simultaneous matches show no such dependence on TGF duration.
- 2 types of WWLLN signal: 1 is the TGF. The other is +IC lightning. Clears up mystery of strongest vs weak discharges associated with TGFS, and why most are simultaneous but outliers exist.

# ARE THERE TWO TYPES OF WWLLN DISCHARGES ASSOCIATED WITH TGFS?

Improvements to WWLLN reveal energy of discharge (Hutchins et al. 2012)



32 WWLLN discharges  $> 0.2$  ms from TGF peak. Mean Energy: 700 J +  
154 WWLLN discharges simultaneous with TGF peak. Mean Energy: 3.1 kJ

Connaughton et al. (2012)

Valerie Connaughton

# THE FUTURE

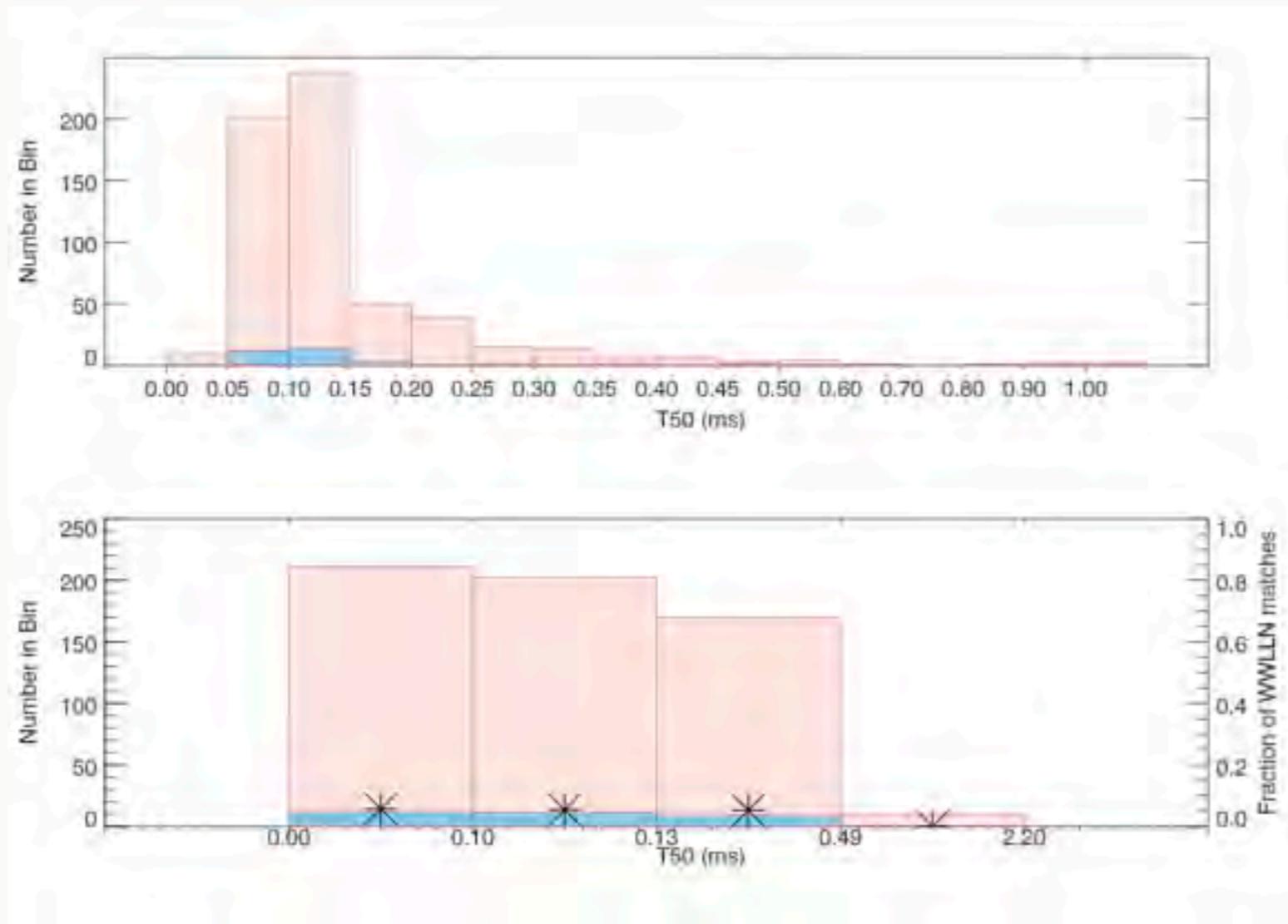
- Over 1000 TGFs from RHESSI and nearing 1000 with GBM. AGILE TGF catalogs. GBM online catalog:<http://gammaray.nsstc.nasa.gov/gbm/science/tgf/>
- Radio observations allow us to study relationship with lightning and to detect currents from electron avalanches, shedding light on TGF gamma-ray production mechanism. We have found 2 types of VLF signal: the TGF itself and +IC lightning from the same flash.
- ATTATT: All-TTE-All-The-Time. Imminent! **Shaolin Xiong (Tuesday)**
- The LAT sees TGFs! And helps us reject CRs. **Eric Grove (Wednesday)**
- Lightning produces gamma rays on the ground! Dwyer et al. (2012), and **poster by Becky Ringuette.**

# BACKUP

- Backup

594 TGFs (triggered + offline)

32 TGFs with a WWLLN discharge  $> 0.2$  ms from peak



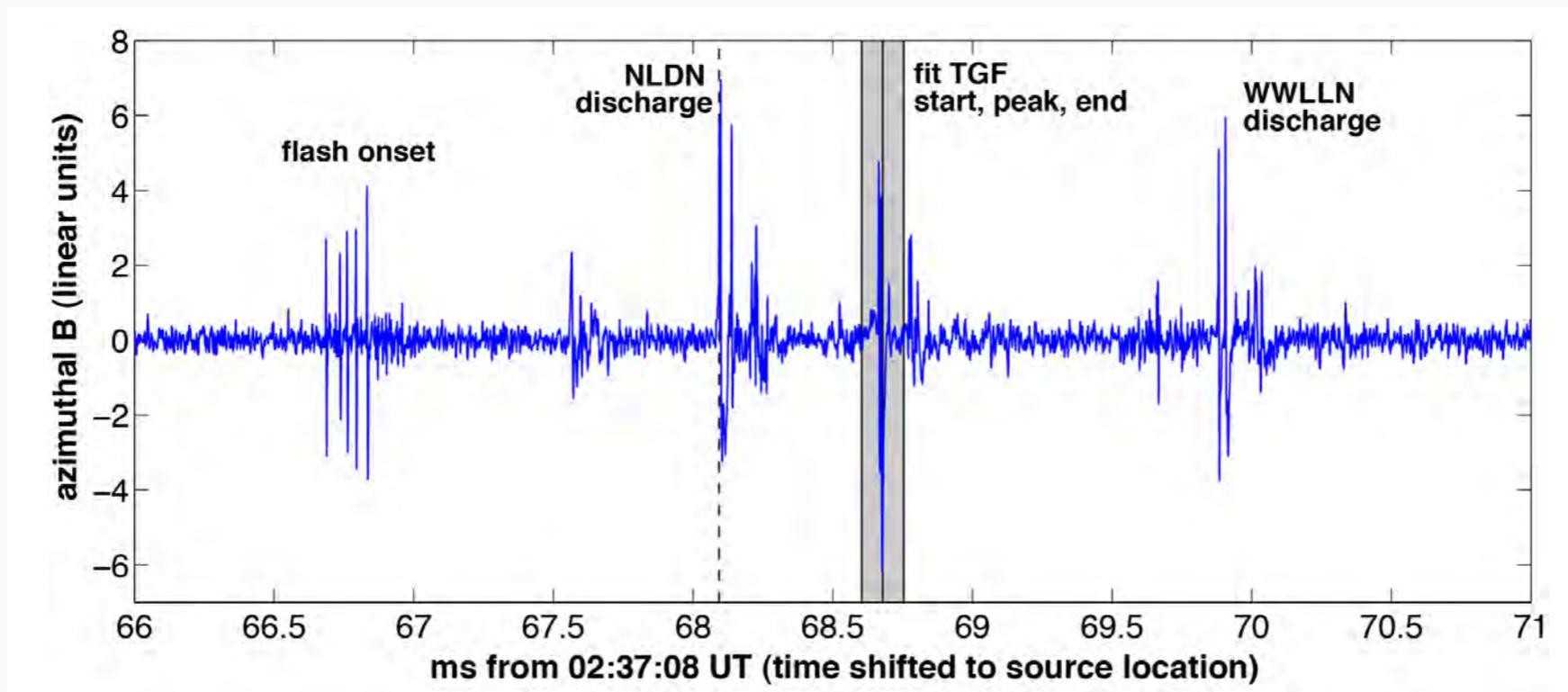
Duration (T50) Distribution

WWLLN Match rate  
as function of  
TGF duration (t50)

# COLLIER AND GREFENSTETTE

- Explain new large catalog, including fainter RHESSI TGFs. Mention weaker ones appear more likely to have a WWLLN match. Why should a TGF property affect detection of associated lightning by WWLLN?

# THE DUKE-GBM TGF IN MORE DETAIL



# GBM-RHESSI T50 DISTRIBUTION

- 170 RHESSI TGFs, 593 GBM TGFs

